# From Hamiltonian to dissipative chaos and back: A primer of active particles





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#### Hamiltonian vs dissipative chaos

- Hamiltonian dynamics: phase volume is conserved, Poincaré recurrence theorem Dissipative dynamics: phase volume shrinks, attractors are observed

with seminal contributions of Hadamard, Birkhoff, KAM, Chirikov and others

(Lorenz, Smale and Williams, Roessler, Hénon, and others)

- Hamiltonian chaos: long history starting from Poincaré treatise on the three-body problem,
- In dissipative systems first examples of transient chaos appeared (van der Pol and van der Mark, Cartright and Levinson, and others) and only later examples of strange attractors

## Hamiltonian dynamics with two degrees of freedom

The simplest Hamiltonian autonomous system where chaos can occur is a particle in a twodimensional potential

One of the first examples: Hénon-Heiles potential

U(x, y) = -

$$H = \frac{p_x^2 + p_y^2}{2} + U(x, y)$$

$$x^2 + y^2 + 2x^2y - \frac{2}{3}y^3$$

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#### Hénon-Heiles example of Hamiltonian chaos



#### Michel Hénon (1931-2013)

#### $\dot{y}$ $\dot{y}$

y = E = 0.12500 0.4 0.3 0.2 0.1 0 -0.1 -0.2 -0.3 -0.4 -0.4 -0.4 -0.3 -0.2 -0.1 0 -0.4 -0.4 -0.3 -0.2 -0.1 0

FIG. 5. Results for E = 0.12500.

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#### The Applicability of the Third Integral Of Motion: Some Numerical Experiments

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## Making Hamiltonian system dissipative

Adding pure dissipation: Energy decreases, attractors are steady states at local minima of potential energy Adding activity:

$$\dot{\vec{r}} = \vec{v} , \qquad \dot{\vec{v}} = -\nabla U + \frac{1}{\mu} (V^2 - v^2) \vec{v}$$

An active velocity-dependent force describes convergence (with rate  $\mu^{-1}$ ) of the speed to the

There are also other models of active particles, in many of them (e.g., in the famous Vicsek model) noise is included

Here I consider deterministic dynamics only

preferred speed V ("cruise control") [e.g., Romanczuk et al., Eur. Phys. J. Spec. Topics, 2012, 202:1-62]



#### Quasiperiodic and chaotic attractors in a harmonic potential

Numerical solution of the equations:

$$\dot{\vec{r}} = \vec{v}$$
,  $\dot{\vec{v}} = -\nabla U + \frac{1}{\mu}(V^2 - v^2)\vec{v}$ ,  $U(x, y) = \frac{x^2 + by^2}{2}$ 

yields quasiperiodic and chaotic attractors, but the latter are rare

Poincaré maps for $b = 2$ and	1.05
different values of activity	
parameter $\mu$ show quasiperiodic	1
(blue) and chaotic (red) attractors	vy



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0.6

### Overactive limit

It is convenient to introduce speed v and direction of motion  $\vec{n}$  via  $\vec{v} = v\vec{n}$ :  $\dot{\vec{r}} = \vec{v}$ ,  $\dot{\vec{v}} = -\nabla U + \frac{1}{\mu}(V^2 - v^2)\vec{v} \Rightarrow$  $\dot{\vec{r}} = v\vec{n}, \quad \dot{v} = \frac{1}{\mu}(V^2 - v^2)v - \mu$ 

Take the overactive limit (very strong cruise control)  $\mu \to 0$ , then v = V and the resulting equations are

$$\dot{\vec{r}} = V\vec{n}$$
  $V\dot{\vec{n}}$ 

In "natural coordinates"  $\vec{n} = (\cos \theta, \sin \theta)$  we obtain

$$\frac{dx}{dt} = V\cos\theta$$
$$\frac{d\theta}{dt} = \frac{1}{V}\left(-\partial_y\right)$$

$$-\nabla U\vec{n}, \quad \dot{\vec{n}} = \frac{-\nabla U + \vec{n}(\nabla U \cdot \vec{n})}{v}$$

$$= -\nabla U + \vec{n}(\nabla U \cdot \vec{n})$$

$$\frac{dy}{dt} = V\sin\theta$$

 $, U\cos\theta + \partial_x U\sin\theta \Big)$ 

#### Hamiltonian formulation

Remarkably, the equations of motion can be formulated as a Hamiltonian system with Hamilton function

$$H(x, y, p_x, p_y) = V\sqrt{p_x^2 + p_y^2} - V^2 \exp\left[-\frac{U(x, y)}{V^2}\right] = 0$$

Because the energy is conserved,  $\sqrt{p_x^2 + p_y^2} = V^{-1} \exp[-U(x, y)V^{-2}]$  and a substitution  $p_x = V^{-1} \exp[-U(x, y)V^{-2}]\cos\theta, p_y = V^{-1} \exp[-U(x, y)V^{-2}]\sin\theta$ leads to the equations in the standard form

The same Hamiltonian describes the ray dynamics in geometrical optics  $H(x, y, p_x, p_y) = \mathbf{1}$ 

$$\sqrt{p_x^2 + p_y^2 - n(x, y)} = 0$$

#### Small velocities: time-scales separation

 $\dot{x} = V \cos \theta, \quad \dot{y} = V \sin \theta,$ 

For small V we have a fast adjustment of the angle heta to the gradient of the potential U according to

$$\dot{\theta} = V^{-1} |\nabla U| \sin(\alpha - \theta), \quad \sin \alpha = -\frac{\partial_y U}{|\nabla U|}, \cos \alpha = -\frac{\partial_x U}{|\nabla U|}$$

2018)]

Near the minimum of the potential, the time-scale separation fails

Close to a potential minimum we can consider the potential as a harmonic one

$$V\dot{\theta} = -\cos\theta\partial_y U + \sin\theta\partial_x U$$

and then slow drift along the gradient toward the minimum of the potential [the dynamics  $\dot{\theta} = \sin(\alpha - \theta)$  was used by Chepizhko & Peruani (PRL, 2013) and Peruani & Aranson (PRL,



#### Motion in a harmonic potential

In a harmonic potential  $U_h = \frac{x^2 + b^2 y^2}{2}$  the velocity parameter can be renormalized to one, and remains only an irrelevant parameter b



Nonuniform distribution on the Poincare map because we work with non-canonical variables: Density ~ exp  $\left| -\frac{U(x, y)}{V^2} \right|$ 

 chaotic motion of particles coming from large values of potential

 quasiperiodic motion of particles starting close to the minimum

![](_page_9_Picture_8.jpeg)

#### Motion in a harmonic potential

![](_page_10_Figure_1.jpeg)

The same dynamics, but the Poincaré map is plotted in canonical variables

### Motion in a harmonic potential: movie

![](_page_11_Figure_1.jpeg)

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#### Chaotic scattering on a potential well

![](_page_12_Figure_1.jpeg)

#### Alignment: dissipative interaction of particles

Aligning coupling of Kuramoto/Viscek type (global)

 $\dot{x}_k = V \cos \theta$ 

$$\dot{\theta}_k = (\epsilon F_y - \partial_y U(\vec{r}_k)) \cos \theta_k - (\epsilon F_x - \partial_x U(\vec{r}_k)) \sin \theta_k \qquad F_x = \sum_1^N \cos \theta_k, \quad F_y = \sum_1^N \sin \theta_k$$

Synchronization: in a harmonic potential particles build a synchronous cluster  $x_1 = \ldots = x_N, y_1 = \ldots = y_N, \theta_1 = \ldots = \theta_N$ , in this cluster the aligning force vanishes

Regularization: the dynamics of the final cluster is Hamiltonian quasiperiodic

$$\theta_k \quad \dot{y}_k = V \sin \theta_k$$

### Alignment: movie

![](_page_14_Figure_1.jpeg)

#### Time-dependent potential and phase volume conservation

I cannot extend the Hamilton function H(x,

time-dependent potential U(x, y, t). Thus, another approach is used - calculation of the phase volume evolution in the full equations

$$\frac{dx}{dt} = V\cos\theta$$
$$\frac{d\theta}{dt} = \frac{1}{V}\left(-\partial_y\right)$$

The divergence rate  $\alpha$  is  $\alpha(t) = W^{-1} \frac{dW}{dt} = \partial_x \dot{x} + \partial_y \dot{y} + \partial_\theta \dot{\theta} =$ 

For a time-independent potential the average divergence rate vanishes  $\langle \alpha(t) \rangle_T = \frac{1}{T} \int_0^T \alpha(t)$ 

$$y, p_x, p_y) = V_v \sqrt{p_x^2 + p_y^2} - \exp\left[-\frac{U(x, y)}{V^2}\right] = 0$$
 t

$$\frac{dy}{dt} = V\sin\theta$$

 $U\cos\theta + \partial_x U\sin\theta$ 

$$= V^{-2}(U_y \dot{y} + U_x \dot{x}) = V^{-2} \left(\frac{dU}{dt} - \frac{\partial U}{\partial t}\right)$$

$$(t')dt' = \frac{U_T - U_0}{V^2 T} \xrightarrow[T \to \infty]{} 0$$

![](_page_15_Picture_13.jpeg)

![](_page_15_Picture_14.jpeg)

#### Time-dependent potential and phase volume conservation

The divergence rate  $\alpha$  is  $\alpha(t) = V^{-2} \left( \frac{dU}{dt} - \frac{\partial U}{\partial t} \right)$ For a time-dependent potential the average divergence rate not necesseraly vanishes Example: breathing potential  $U(x, y, t) = \frac{a(1 + \Gamma \cos(\omega t))x^2 + b(1 + \Gamma \sin(\omega t))y^2}{2}$ 

![](_page_16_Figure_5.jpeg)

![](_page_16_Picture_6.jpeg)

![](_page_16_Picture_7.jpeg)

### Interacting particles

I cannot write a Hamiltonian formulation for particles interacting with some potential  $U(\vec{r}_1, \vec{r}_2)$ , but one can check the phase volume conservation  $\dot{x}_{1,2} = V_{1,2} \cos \theta_{1,2} ,$  $\dot{y}_{1,2} = V_{1,2} \sin \theta_{1,2}$ ,  $\dot{\theta}_{1,2} = \frac{1}{M_{1,2}V_{1,2}} \left( -\frac{\partial U}{\partial y_{1,2}} \right)$ 

In general, masses M and velocities V for two particles are different and the phase volume divergence is

$$\alpha(t) = \sum_{m=1,2} \left( \frac{\partial \dot{x}_m}{\partial x_m} + \frac{\partial \dot{y}_m}{\partial y_m} + \frac{\partial \dot{\theta}_m}{\partial \theta_m} \right) = \sum_{m=1,2} (M_m V_m^2)^{-1} \left( \dot{x}_m \frac{\partial U}{\partial x_m} + \dot{y}_m \frac{\partial U}{\partial y_m} \right)$$

$$\frac{\partial U}{\partial x_{1,2}} \cos \theta_{1,2} + \frac{\partial U}{\partial x_{1,2}} \sin \theta_{1,2} \right) .$$

![](_page_17_Picture_7.jpeg)

#### Identical and non-identical interacting particles

$$\alpha(t) = \sum_{m=1,2} \left( \frac{\partial \dot{x}_m}{\partial x_m} + \frac{\partial \dot{y}_m}{\partial y_m} + \frac{\partial \dot{\theta}_m}{\partial \theta_m} \right) = \sum_{m=1,2} \left( M_m V_m^2 \right)^{-1} \left( \dot{x}_m \frac{\partial U}{\partial x_m} + \dot{y}_m \frac{\partial U}{\partial y_m} \right)$$

 $\alpha(t) = (MV^2)^{-1} \frac{dU}{dt}$  and on a long term, the phase volume is conserved on average

If the particles are different  $M_1V_1^2 \neq M_2V_2^2$ , the phase volume is not conserved

If particles are identical, then the divergence rate is the total derivative of the potential

#### Many non-identical interacting particles are dissipative

We simulated several particles interacting via a smooth repulsing potential

 $U_{ij}(R) = \begin{cases} D | (R/\sigma)^2 - 1 |^7 & R < \sigma , \\ 0 & R \ge \sigma , \end{cases}$ in a harmonic confining potential

![](_page_19_Figure_3.jpeg)

FIG. 5. Particles with potential interaction (parameters  $\sigma =$ 1,  $D = 10^4$ , time of averaging  $10^5$ .) Red: rate vs  $\delta_M$  for  $\delta_V = 0$  (the masses of particles are  $1 \pm \delta_M/2$ ); blue: rate vs  $\delta_V$  (the velocities of particles are  $V = 1 \pm \delta_V/2$ ) for  $\delta_M = 0$ . Squares: two particles, circles: five particles; triangles: 10 particles. Lines: fits according to the square rate  $\sim \delta^2$ . All the rates are scaled by the particle number (i.e., convergence) rate per particle).

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![](_page_19_Picture_6.jpeg)

#### Conservative vs dissipative dynamics

Classical particles

Overactive particle ina timeindependent potential

Synchronized cluster of aligned particles

Interacting identical overactive particles

Active particles

Alignment of many overactive particles

Particle in a time-dependent potential

Interacting non-identical overactive particles

![](_page_20_Picture_10.jpeg)

![](_page_20_Figure_11.jpeg)